



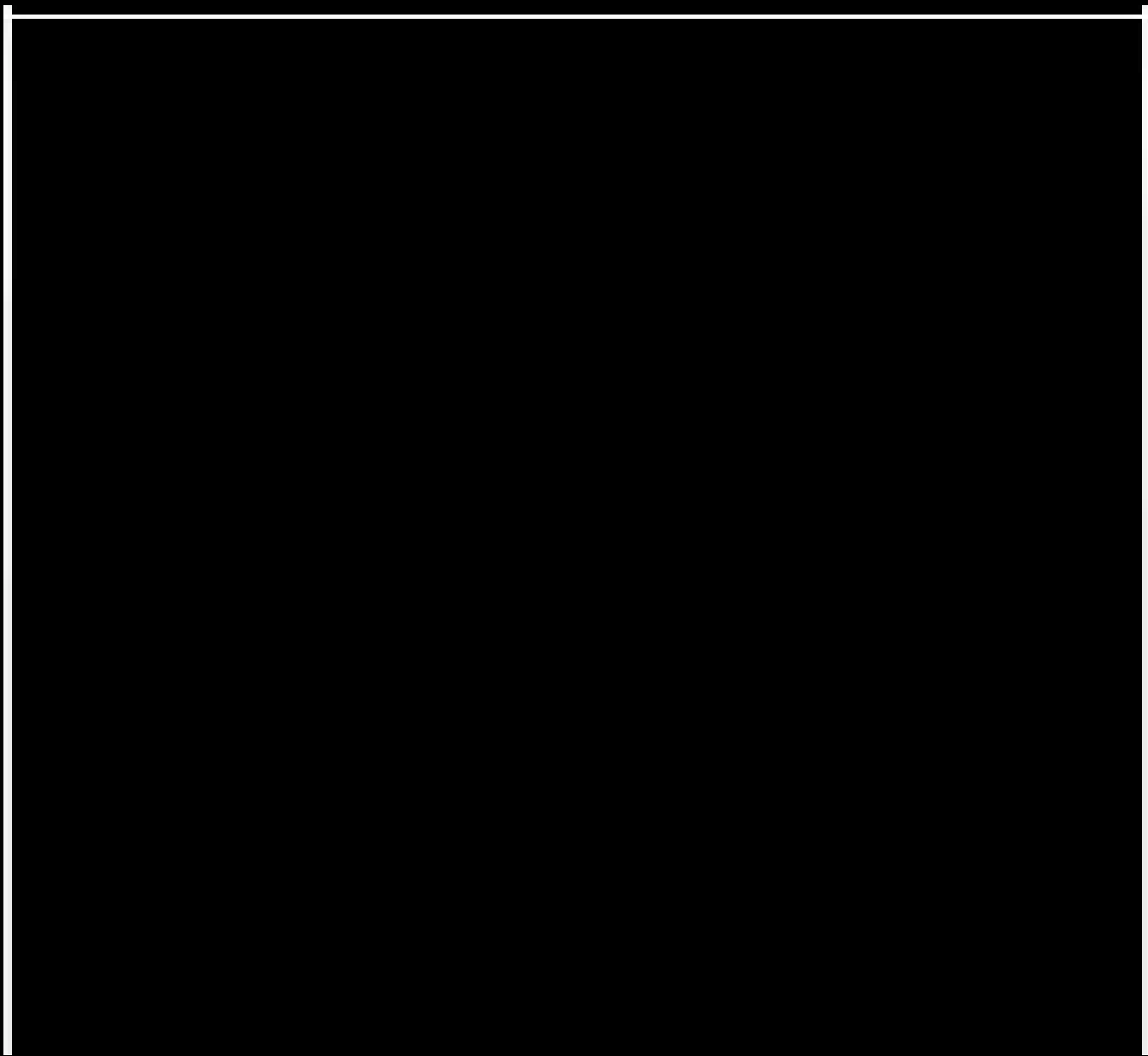
Terminal Guidance Navigation for an Asteroid Impactor Spacecraft

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Introduction

- On July 4, 2005, the Deep Impact impactor spacecraft successfully collided with comet 9P/Tempel 1, while the main spacecraft flew by and shuttered images which captured the impact
 - 1st hypervelocity impact of a primitive Solar System body
 - Not primary goal of mission, but it did demonstrate that such an impact could be accomplished with current technologies and relatively modest budget
- For relatively small asteroids and short turnaround times from detection to impact, kinetic energy technique recommended as the most practical and cost effective technique for deflection



AutoNav

- DI impact made possible by onboard closed-loop autonomous navigation system (AutoNav) for the terminal guidance
- AutoNav originally developed as a technology demonstration on Deep Space 1
- To date, 5 missions, using 4 different spacecraft, have used AutoNav, primarily to impact a comet nucleus (DI), and to track asteroid or comet nucleus through closest approach for a flyby
 - Deep Space 1 (cruise and flyby of comet Borrelly)
 - Stardust (flyby of asteroid Annefrank and comet Wild 2)
 - Deep Impact (Impactor and Flyby spacecraft imaging for comet Tempel 1)
 - EPOXI (flyby of comet Hartley 2)
 - Stardust NExT (flyby of comet Tempel 1)
- Technology for tracking nucleus through flyby identical to that needed for closed loop control for flyby

Comparison of AutoNav Approach Parameters for Past Missions and Potential Impactor Mission

Mission/Target	Flyby Radius (km)	Flyby Velocity (km/s)	Approach Phase (deg)	Target size (km)
DS1/Borrelly	2171	16.6	65	4.8
STARDUST/Annefrank	3076	7.2	150	4.8
STARDUST/Wild 2	237	6.1	72	4.0
DI/Tempel 1	500/0	10.2	62	6.0
EPOXI/Hartley 2	694	12.3	86	1.6
STARDUST NExT/Tempel 1	182	10.9	82	6.0
Potential KI Scenarios	0	~3 to 20	~60 - 140	~0.100 – 0.300

Brief Background on Deep Space Navigation

- Step 1: design trajectory to intercept asteroid to satisfy mission constraints
 - Launch vehicle
 - Delivered mass
 - Fuel required
 - Approach velocity, phase angle
- Step 2: navigate reference trajectory from launch to impact
 - Standard techniques of ground-based navigation used for launch, cruise, and early approach, using primarily radiometric tracking data
 - Ground navigation delivers spacecraft to interface location at predetermined time before impact, at which time light-time delays require that onboard control take over
 - For previous missions, this point ranged between 30 min and 2 hours prior to closest approach/impact

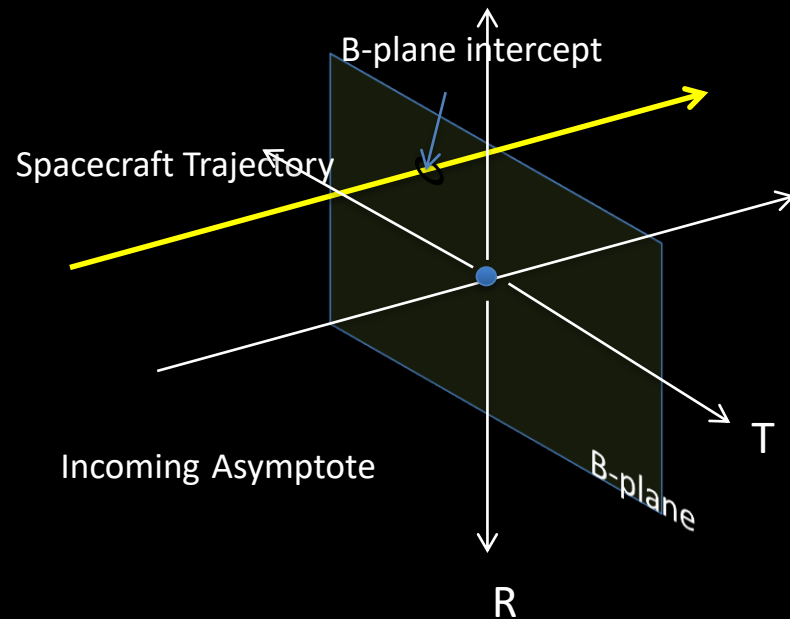
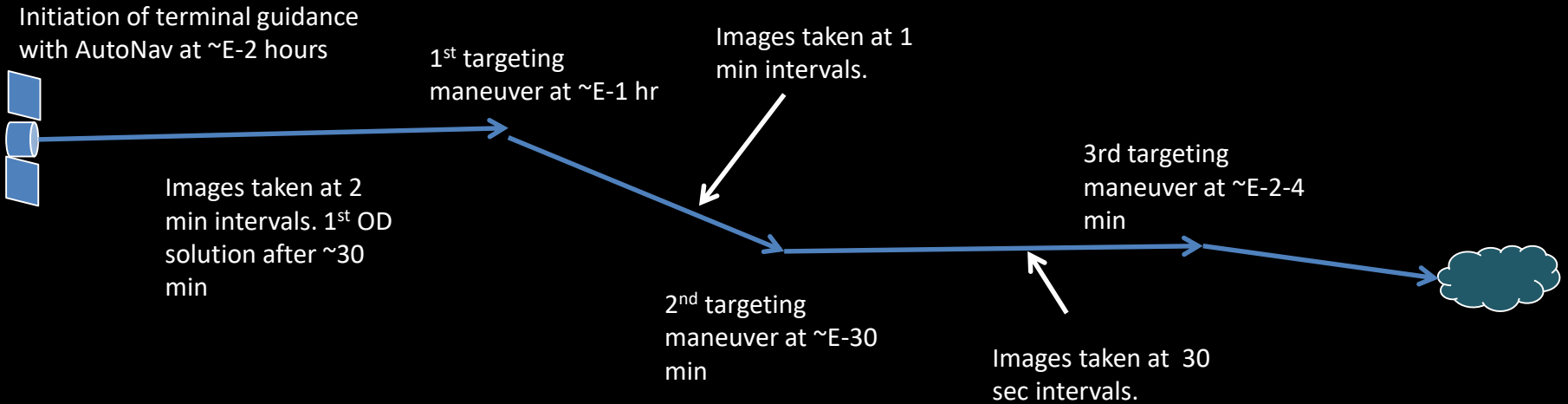
AutoNav Description

- Entirely self-contained system uses onboard camera to take images of target body to compute target relative navigation solution
 - Does not require radio link to other s/c or the Earth
- 3 main components of AutoNav
 - Image processing element to extract target center-of-figure information
 - Orbit determination element to combine set of target centroid information in least-squares filter estimate of s/c trajectory
 - Maneuver planning and execution element to compute delta-V needed to hit target
 - For Deep Impact, 3 maneuvers were used for targeting. These were placed at E-90 min, E-35 min, and E-12.5 min

Some Notes on Onboard Orbit Estimation

- Different from “follow the target” techniques using PN guidance in that the complete relative orbit solution is estimated
- AutoNav needed to be robust against various failure modes
 - Failure of camera during any portion of terminal guidance
 - Loss of images due to corruption from stray light, cosmic rays, etc.
 - Large variations in COB due to shape/phase effects
 - Attitude disruptions due to particle impacts (for comets)
 - Unknown size, brightness, shape, orientation of target object
 - Failure of TCM to execute

Notional Targeting Scenario for KI



Optical Navigation (Opnav)

- Opnav images provided by onboard camera
- Provides only target-relative navigation information (ground-based radiometric data provides Earth-referenced navigation information)
- Key parameters for camera include IFOV (angular resolution of single pixel), sensitivity
 - These, along with V_{inf} and approach phase, determine when target object becomes resolved, accuracy of measurement, earliest detection, and ability to see stars along with object to provide inertial reference for observations
- Note that for the deflection scenarios we are examining, the target object will almost always be unresolved at start of terminal guidance, and may remain so until < 5 minutes to impact
- Prior observations of target body, either by orbiting or earlier flyby, can dramatically improve Opnav performance since characteristics of body (shape, size, orientation) will be known

Attitude Knowledge

- Errors in attitude knowledge directly affects accuracy of OD
 - Must estimate attitude error as part of filter which degrades strength of target relative angular information
- “Stellar mode” attitude knowledge
 - Stars available in navigation camera, attitude knowledge near perfect
- Star tracker/IMU
 - Degraded attitude knowledge depending on Star tracker/IMU information
 - Past experience suggests using IMU propagation only
 - IMU bias and drift primary source of error for terminal guidance because these can be difficult to separate from translational motion
 - 2 general classes of IMU capability needed for KI (MIMU, SSIRU)
 - Less capable IMUs (e.g., LN 200) not good enough for this purpose

Monte Carlo Simulations

- Impactor targeting accuracy assessed through Monte Carlo simulations
 - Confidence in simulations have increased with comparison against actual flight performance
 - Thus far, flight performance has fallen within envelope of simulated cases
- Previous studies* looked at KI performance across a range of scenarios
- Targeted trade studies can be done for specific missions to assess performance for a range of parameters
 - Camera specifications (focal length, sensitivity, frame rate)
 - Image processing settings
 - Spacecraft attitude control/knowledge requirements
 - Spacecraft thruster performance
 - Orbit determination filter tuning
- A quick set of simulations were run for the PDC 2017 KI scenario
 - Results very preliminary – no time to tune filter, adjust parameters, etc.

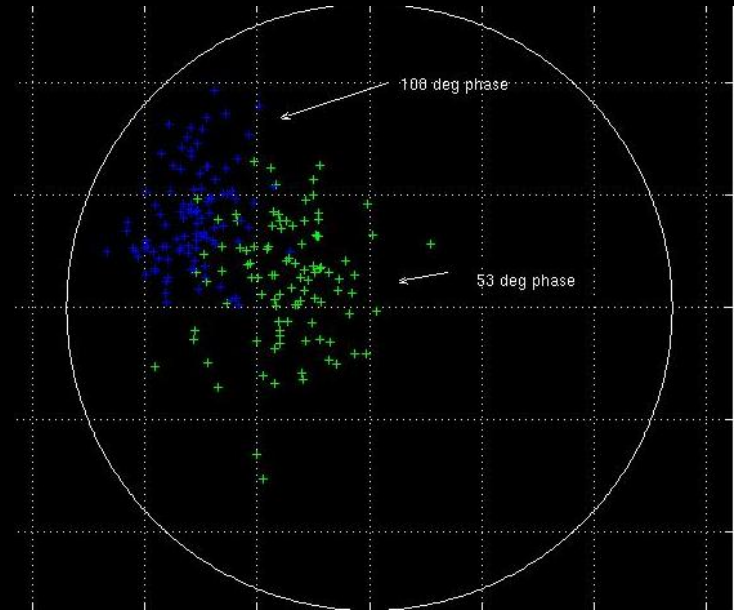
* Bhaskaran, S., Kennedy, B., “Closed loop terminal guidance navigation for a kinetic impactor spacecraft”, *Acta Astronautica*, 103, (2014), pp. 322-332

Simulation Results (from Bhaskaran, Kennedy, 2014 study)

Case	Vinf (km/s)	Phase angle (deg)	Stellar reference		SSIRU	
			100 m	300 m	100 m	300 m
1	7.5	30	98.8%	100.0%	85.5%	100.0%
2	7.5	80	96.5%	100.0%	73.8%	99.2%
3	12.5	140	56.6%	99.4%	53.8%	90.6%
4	20	5	100.0%	100.0%	75.4%	99.6%

Simulation Results (2017 PDC Scenario Case)

- Two cases looked at
 - $V_{\text{inf}} = 15 \text{ km/s}$, Phase angle = 53 deg
 - $V_{\text{inf}} = 12.7 \text{ km/s}$, Phase angle = 108 deg
- 3 attitude modes
 - Stellar reference (near perfect attitude knowledge)
 - SSIRU class gyro
 - MIMU class gyro



Case	V_{inf} (km/s)	Phase angle (deg)	Stellar reference	SSIRU	MIMU
1	15.0	53	100%	100%	38%
2	12.7	108	100%	100%	38%

Conclusion

- Attitude knowledge mode is the single biggest factor in determining impact success
 - With stellar reference, probability of success fairly high
 - Otherwise, must have very stable IMU
- Phase angle second largest effect
 - High value in designing reference trajectories which lower approach phase angle
- Precursor mission valuable for increasing chance of success
 - Can correct for COB, phase angle effects
 - Minimize sizes of maneuvers needed to remove larger target body ephemeris errors

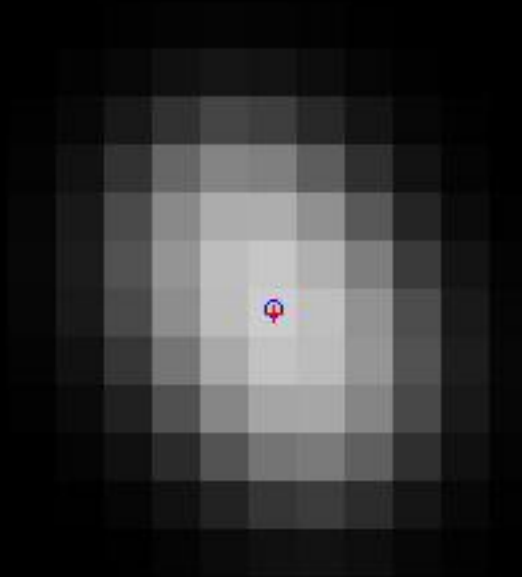
Backup

Simulation Parameters

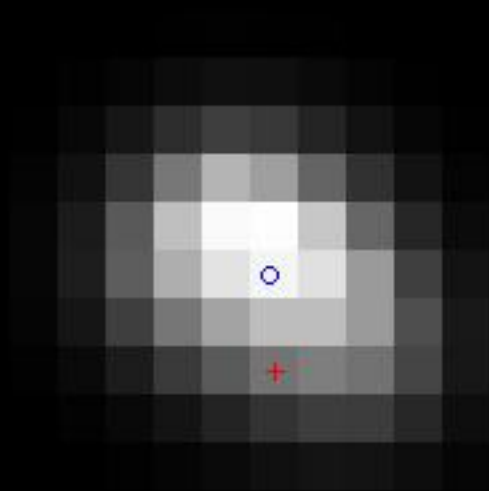
Initial asteroid-relative state error	Position: 30 km Velocity: 5 cm/s
Gates model maneuver execution error	Fixed magnitude: 4.3 mm/s Proportional magnitude: 10% Fixed direction: 4 mm/s Proportional direction: 3.1%
Gyro errors (MIMU class)	Rate bias: 0.005 deg Angle random walk: 0.005 deg/sqrt(hr)
Gyro errors (SSIRU class)	Rate bias: 0.0005 deg Angle random walk: 0.0005 deg/sqrt(hr)
Asteroid size	130 x 90 x 90 m 390 x 260 x 260 m
Asteroid pole orientation	RA: 0 to 360 deg, uniform Dec: -90 to 90 deg, uniform

All errors values are 1 sigma unless otherwise noted

Example of Phase Effects in Final Image



Phase = 5 deg



Phase = 80 deg



Phase = 140 deg